#### Notes

- Assignment 4 due "today" (when I check email tomorrow morning)
- Don't be afraid to make assumptions, approximate quantities, ...
  - In particular, method for computing time step bound (look at max eigenvalue of Jacobian) already made lots of assumptions about linearity etc. so it won't hurt to make a few more!

## **Fluid mechanics**

• We already figured out the equations of motion for continuum mechanics

 $\rho \ddot{x} = \nabla \cdot \sigma + \rho g$ 

· Just need a constitutive model

$$\sigma = \sigma(x, t, \varepsilon, \dot{\varepsilon})$$

- We'll look at the constitutive model for "Newtonian" fluids today
  - Remarkably good model for water, air, and many other simple fluids
  - Only starts to break down in extreme situations, or more complex fluids (e.g. viscoelastic substances)

## **Inviscid Euler model**

- Inviscid=no viscosity
- · Great model for most situations
  - Numerical methods end up with viscosity-like error terms anyways...
- Constitutive law is very simple:

$$\sigma_{ij} = -p\delta_{ij}$$

- New scalar unknown: pressure p
- Barotropic flows: p is just a function of density (e.g. perfect gas law p=k(ρ-ρ<sub>0</sub>)+p<sub>0</sub> perhaps)
- For more complex flows need heavy-duty thermodynamics: an equation of state for pressure, equation for evolution of internal energy (heat), ...

# Lagrangian viewpoint

- We've been working with Lagrangian methods so far
  - Identify chunks of material, track their motion in time, differentiate world-space position or velocity w.r.t. material coordinates to get forces
  - In particular, use a mesh connecting particles to approximate derivatives (with FVM or FEM)
- · Bad idea for most fluids
  - [vortices, turbulence]
  - At least with a fixed mesh...

## **Eulerian viewpoint**

- Take a fixed grid in world space, track how velocity changes at a point
- Even for the craziest of flows, our grid is always nice
- (Usually) forget about object space and where a chunk of material originally came from
  - Irrelevant for extreme inelasticity
  - Just keep track of velocity, density, and whatever else is needed

## **Conservation laws**

- Identify any fixed volume of space
- Integrate some conserved quantity in it (e.g. mass, momentum, energy, ...)
- Integral changes in time only according to how fast it is being transferred from/to surrounding space
  - Called the flux
  - [divergence form]  $\frac{\partial}{\partial t} \int_{\Omega} q = -\int_{\partial \Omega} f(q) \cdot n$

 $q_t + \nabla \cdot f = 0$ 

# Conservation of Mass

- Also called the continuity equation (makes sure matter is continuous)
- Let's look at the total mass of a volume (integral of density)
- Mass can only be transferred by moving it: flux must be  $\rho \textbf{u}$

$$\frac{\partial}{\partial t} \int_{\Omega} \rho = -\int_{\partial \Omega} \rho u \cdot m$$
$$\rho_t + \nabla \cdot (\rho u) = 0$$

# **Material derivative**

- A lot of physics just naturally happens in the Lagrangian viewpoint
  - E.g. the acceleration of a material point results from the sum of forces on it
  - How do we relate that to rate of change of velocity measured at a fixed point in space?
  - Can't directly: need to get at Lagrangian stuff somehow
- The material derivative of a property q of the material (i.e. a quantity that gets carried along with the fluid) is

# $\frac{Dq}{Dt}$

### Finding the material derivative

 Using object-space coordinates p and map x=X(p) to world-space, then material derivative is just

$$\frac{D}{Dt}q(t,x) = \frac{d}{dt}q(t,X(t,p))$$
$$= \frac{\partial q}{\partial t} + \nabla q \cdot \frac{\partial x}{\partial t}$$
$$= q_t + u \cdot \nabla q$$

 Notation: u is velocity (in fluids, usually use u but occasionally v or V, and components of the velocity vector are sometimes u,v,w)

#### **Compressible Flow**

- In general, density changes as fluid compresses or expands
- When is this important?
  - Sound waves (and/or high speed flow where motion is getting close to speed of sound - Mach numbers above 0.3?)
  - Shock waves
- Often not important scientifically, almost never visually significant
  - Though the **effect** of e.g. a blast wave is visible! But the shock dynamics usually can be hugely simplified for graphics

## **Incompressible flow**

- So we'll just look at incompressible flow, where density of a chunk of fluid never changes
  - Note: fluid density may not be constant throughout space different fluids mixed together...
- That is, Dρ/Dt=0

# Simplifying

- Incompressibility:  $\frac{D\rho}{Dt} = \rho_t + u \cdot \nabla \rho = 0$
- Conservation of mass:  $\rho_t + \nabla \cdot (\rho u) = 0$

$$\rho_t + \nabla \rho \cdot u + \rho \nabla \cdot u = 0$$

- Subtract the two equations, divide by  $\rho$ :

 $\nabla \cdot \boldsymbol{u} = \boldsymbol{0}$ 

- Incompressible == divergence-free velocity
  - Even if density isn't uniform!

#### **Conservation of momentum**

• Short cut - in  $\rho \ddot{x} = \nabla \cdot \sigma + \rho g$ use material derivative:

$$\rho \frac{Du}{Dt} = \nabla \cdot \sigma + \rho g$$
$$\rho(u_t + u \cdot \nabla u) = \nabla \cdot \sigma + \rho g$$

- Or go by conservation law, with the flux due to transport of momentum and due to stress:
  - Then simplify a bit using conservation of mass  $(\rho u)_t + \nabla \cdot (u\rho u \sigma) = \rho g$

#### Inviscid momentum equation

 Plug in simplest consitutive law (σ=-pδ) from before to get

$$\rho(u_t + u \cdot \nabla u) = -\nabla p + \rho g$$
$$u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g$$

 Together with conservation of mass: the Euler equations

### **Incompressible inviscid flow**

• So the equations are:  $u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g$ 

 $\nabla \cdot u = 0$ 

- 4 equations, 4 unknowns (u, p)
- Pressure p is just whatever it takes to make velocity divergence-free
- In fact, incompressibility is a hard constraint; div and grad are transposes of each other and pressure p is the Lagrange multiplier
  - Just like we figured out constraint forces before...

#### **Pressure solve**

• To see what pressure is, take divergence of momentum equation

 $\nabla \cdot \left( u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p \right) = 0$ 

$$\nabla \cdot \left(\frac{1}{\rho} \nabla p\right) = -\nabla \cdot \left(u_t + u \cdot \nabla u\right)$$

- For constant density, just get Laplacian (and this is Poisson's equation)
- Important numerical methods use this approach to find pressure

## Projection

• Note that  $\nabla \cdot u_t = 0$  so in fact

 $\nabla \cdot \frac{1}{\rho} \nabla p = -\nabla \cdot \left( u \cdot \nabla u \right)$ 

- After we add 
  <sup>v</sup>p/ρ to u<sup>•</sup>νu, divergence must be zero
- So if we tried to solve for additional pressure, we get zero
- Pressure solve is linear too
- Thus what we're really doing is a projection of u•∇u onto the subspace of divergence-free functions: u<sub>t</sub>+P(u•∇u)=g

# Viscosity

- In reality, nearby molecules travelling at different velocities occasionally bump into each other, transferring energy
  - Differences in velocity reduced (damping)
  - Measure this by strain rate (time derivative of strain, or how far velocity field is from rigid motion)
  - · Add terms to our constitutive law

### Strain rate

- At any instant in time, measure how fast chunk of material is deforming from its current state
  - Not from its original state
  - So we're looking at infinitesimal, incremental strain updates
  - Can use linear Cauchy strain!
  - (In fact, in solids, this leads to a more advanced "true" strain arrived at by integrating infinitesimal strain increments... but not important here)
- So the strain rate tensor is  $\dot{\varepsilon}_{ij} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$

#### **Viscous stress**

• As with linear elasticity, end up with two parameters if we want isotropy:

 $\sigma_{ij}^{viscous} = 2\mu \dot{\varepsilon}_{ij} + \lambda \dot{\varepsilon}_{kk} \delta_{ij}$ 

- $\mu$  and  $\lambda$  are coefficients of viscosity (first and second)
- These are not the Lame coefficients! Just use the same symbols
- +  $\lambda$  damps only compression/expansion
- Usually  $\lambda \approx -2/3\mu$  (exact for monatomic gases)
- So end up with  $\sigma_{ij}^{viscous} = \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \frac{2}{3} \frac{\partial u_k}{\partial x_k} \delta_{ij} \right)$

#### **Navier-Stokes**

- Navier-Stokes equations include the viscous stress
- Incompressible version:

$$u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g + \frac{1}{\rho} \nabla \cdot \mu \Big( \nabla u + \nabla u^T \Big)$$
$$\nabla \cdot u = 0$$

 Often (but **not** always) viscosity μ is constant, and this reduces to

$$u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g + \frac{\mu}{\rho} \nabla^2 u$$

### **Inviscid boundaries**

- · Basic choice:
  - At a closed boundary ("wall") use the "no-stick" condition  $u \cdot n = 0$
  - Or if the boundary is moving in the normal direction,  $u \cdot n = u_{wall} \cdot n$
  - At an open boundary ("free surface") specify the traction, i.e. the pressure. For example:

$$p = p_{atm} \quad \frac{\partial u}{\partial n} = 0$$

• Condition on normal derivative of u implicit

## **Boundary conditions**

- We'll sidestep this issue until it comes up in numerical methods
- There are some subtle mathematical details (and open problems) relating to what exactly you can or need to specify
- Generally: specify some mix of velocity and traction at the boundary
  - Depends on whether or not you have viscosity

#### **Viscous (friction) boundaries**

- Can use "no-slip" condition on walls: u=0
  - Or u=u<sub>wall</sub>
- Traction is no longer so simple at a free surface (stress tensor has more than just pressure)

# Other quantities

- We may want to carry around auxiliary quantities
  - E.g. temperature, the type of fluid (if we have a mix), concentration of smoke, etc.
- Use material derivative as before
- E.g. if quantity doesn't change, just is transported ("advected") around:

$$\frac{Dq}{Dt} = q_t + \underbrace{u \cdot \nabla q}_{advection} = 0$$

## **Pressure stuff**

- Just as we solved for pressure before, can do the same here
- But we take the divergence of the viscosity term as well

#### What now?

- · Can solve numerically the full equations
  - Will do this later
  - Not so simple, could be expensive (3D)
- Or make assumptions and simplify them, then solve numerically
  - Simplify flow (e.g. irrotational)
  - Simplify dimensionality (e.g. go to 2D)

#### **Potential flow**

• If velocity is irrotational:

 $\nabla \times u = 0$ 

- Which it often is in simple laminar flow
- Then there must be a stream function (potential) such that:  $u = \nabla \phi$
- Solve for incompressibility:  $\nabla \cdot \nabla \phi = 0$