## CPSC 425: Computer Vision



Lecture 2: Image Formation
( unless otherwise stated slides are taken or adopted from Bob Woodham, Jim Little and Fred Tung )

## Menu for Today (January 11, 2022)

## Topics:

- Image Formation
- Projection
- Cameras and Lenses


## Redings:

- Today's Lecture: Forsyth \& Ponce (2nd ed.) 1.1.1 - 1.1.3
- Next Lecture: Forsyth \& Ponce (2nd ed.) 4.1, 4.5


## Reminders:

- Complete Assignment 0 (ungraded) by Monday, January 1
- Please sign up for Piazza (116 students signed up so far)
- CoLab and Jupyter Notebooks for assignments

Today’s "fun" Example

## Today’s "fun" Example



Photo credit: reddit user Liammm

## Today’s "fun" Example: Eye Sink Illusion

## 



## Salvador Dali - Pareidolia



## Lecture 1: Re-cap

Types of computer vision problems:

- Computing properties of the 3D world from visual data (measurement)
- Recognition of objects and scenes (perception and interpretation)
- Search and interact with visual data (search and organization)
- Manipulation or creation of image or video content (visual imagination)

Computer vision challenges:

- Fundamentally ill-posed
- Enormous computation and scale
- Lack of fundamental understanding of how human perception works


## Lecture 1: Re-cap

Computer vision technologies have moved from research labs into commercial products and services. Examples cited include:

- broadcast television sports
- electronic games (Microsoft Kinect)
- biometrics
- image search
- visual special effects
- medical imaging
- robotics
... many others


## Lecture 2: Goal

## To understand how images are formed

(and develop relevant mathematical concepts and abstractions)

## What is Computer Vision?

Compute vision, broadly speaking, is a research field aimed to enable computers to process and interpret visual data, as sighted humans can.

Sensing Device Interpreting Device



## What is Computer Vision?

Compute vision, broadly speaking, is a research field aimed to enable computers to process and interpret visual data, as sighted humans can.


## Overview: Image Formation, Cameras and Lenses

The image formation process that produces a particular image depends on

- Lightening condition
- Scene geometry
- Surface properties
- Camera optics and viewpoint
source


Sensor (or eye) captures amount of light reflected from the object

## Light and Color: A Short Preview



Visible light is electromagnetic radiation in the 400nm-700nm band of wavelengths

## Light and Color: A Short Preview



Visible light is electromagnetic radiation in the 400nm-700nm band of wavelengths

- Black is the absence of light
- Sunlight is a spectrum of light


## Light and Color: A Short Preview



Visible light is electromagnetic radiation in the 400nm-700nm band of wavelengths


## Light and Color: A Short Preview



Visible light is electromagnetic radiation in the 400nm-700nm band of wavelengths


Light also behaves as particles with specific wavelengths

- photons; that travel in straight lines within a medium


## Overview: Image Formation, Cameras and Lenses

The image formation process that produces a particular image depends on

- Lightening condition
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- Surface properties
- Camera optics and viewpoint
source


Sensor (or eye) captures amount of light reflected from the object

## (small) Graphics Review



## (small) Graphics Review



## (small) Graphics Review



## (small) Graphics Review

Surface reflection depends on both the viewing $\left(\theta_{v}, \phi_{v}\right)$ and illumination $\left(\theta_{i}, \phi_{i}\right)$ direction, with Bidirectional Reflection Distribution Function: $\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)$


## (small) Graphics Review

Surface reflection depends on both the viewing $\left(\theta_{v}, \phi_{v}\right)$ and illumination $\left(\theta_{i}, \phi_{i}\right)$ direction, with Bidirectional Reflection Distribution Function: $\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)$


## (small) Graphics Review

Surface reflection depends on both the viewing ( $\theta_{v}, \phi_{v}$ ) and illumination $\left(\theta_{i}, \phi_{i}\right)$ direction, with Bidirectional Reflection Distribution Function: $\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)$


## (small) Graphics Review

Surface reflection depends on both the viewing $\left(\theta_{v}, \phi_{v}\right)$ and illumination $\left(\theta_{i}, \phi_{i}\right)$ direction, with Bidirectional Reflection Distribution Function: $\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)$
source

Lambertian surface:

$\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)=\frac{\rho_{d}}{\pi}$

$$
L=\frac{\rho_{d}}{\pi} I(\vec{i} \cdot \vec{n})
$$



## (small) Graphics Review

Question: What are the simplifying assumptions we are making here?
source

Lambertian surface:

$\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)=\frac{\rho_{d}}{\pi}$

$$
L=\frac{\rho_{d}}{\pi} I(\vec{i} \cdot \vec{n})
$$



Slide adopted from: Ioannis (Yannis) Gkioulekas (CMU)

## (small) Graphics Review

Question: What are the simplifying assumptions we are making here?

1. BRDF is the same everywhere (i.e., surface has identical properties everywhere)
source

Lambertian surface:
2. Light spectra is absorbed uniformly by the surface (no change in color)

$$
\begin{aligned}
& \operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)=\frac{\rho_{d}}{\pi} \\
& L=\frac{\rho_{d}}{\pi} I(\vec{i} \cdot \vec{n})
\end{aligned}
$$

## (small) Graphics Review

Surface reflection depends on both the viewing $\left(\theta_{v}, \phi_{v}\right)$ and illumination $\left(\theta_{i}, \phi_{i}\right)$ direction, with Bidirectional Reflection Distribution Function: $\operatorname{BRDF}\left(\theta_{i}, \phi_{i}, \theta_{v}, \phi_{v}\right)$

Lambertian surface:


Mirror surface: all incident light reflected in one directions $\left(\theta_{v}, \phi_{v}\right)=\left(\theta_{r}, \phi_{r}\right)$

## Cameras

Old school film camera
Digital CCD/CMOS camera


## Cameras

Old school film camera
Digital CCD/CMOS camera


## Let's say we have a sensor ...

Digital CCD/CMOS camera


## Let's say we have a sensor ...

Digital CCD/CMOS camera


## Let's say we have a sensor ...

Digital CCD/CMOS camera

digital sensor (CCD or CMOS)

## ... and the object we would like to photograph

What would an image taken like this look like?


## Bare-sensor imaging



## Bare-sensor imaging



## Bare-sensor imaging



## Bare-sensor imaging



All scene points contribute to all sensor pixels

## Bare-sensor imaging

## All scene points contribute to all sensor pixels

## Pinhole Camera



What would an image taken like this look like?

## Pinhole Camera



Slide Credit: Ioannis (Yannis) Gkioulekas (CMU)

## Pinhole Camera



Each scene point contributes to only one sensor pixel

## Camera Obscura (latin for "dark chamber")

illum in tabula per radios Solis, quam in coelo contingit: hoc eft,fil in ccelo fuperior pars deliquiū patiatur, in radiis apparebit inferior deficere, vt ratio exigit optica.


Sic nos exactè Anno.1544. Louanii celipfim Solis obferuauimus, inuenimusq́; deficere paulò plus $\not \approx$ dex-

Reinerus Gemma-Frisius observed an eclipse of the sun at Louvain on January 24, 1544. He used this illustration in his book, "De Radio Astronomica et Geometrica," 1545. It is thought to be the first published illustration of a camera obscura.

## Camera Obscura (latin for "dark chamber")

```
illum in tabula per radios Solis, quamm in coelo contin- git: hoc eft,fil in ccelo fuperior pars deliquiū patiatur, in radiis apparebit inferior deficere,vt ratio exigit optica.
Soles delignuinm Amo Chirin
154.4. Dio 24. Januarí
Conami
principles behind the pinhole camera or camera obscura were first mentioned by Chinese philosopher Mozi (Mo-Ti) (470 to 390 BCE)
```



Sic nos exactè Anno.1544. Louanii celipfim Solis obferuauimus, inuenimuśq; deficere paulò plus $\underset{q}{\text { q. }}$ dex-

Reinerus Gemma-Frisius observed an eclipse of the sun at Louvain on January 24, 1544. He used this illustration in his book, "De Radio Astronomica et Geometrica," 1545. It is thought to be the first published illustration of a camera obscura.

## First Photograph on Record

La table servie


## Pinhole Camera

A pinhole camera is a box with a small hall (aperture) in it


Forsyth \& Ponce (2nd ed.) Figure 1.2

## Pinhole Camera

A pinhole camera is a box with a small hall (aperture) in it


Forsyth \& Ponce (2nd ed.) Figure 1.2

## Image Formation



Forsyth \& Ponce (2nd ed.) Figure 1.1

## Accidental Pinhole Camera



## Pinhole Camera (Simplified)

f' is the focal length of the camera


## Pinhole Camera (Simplified)

f' is the focal length of the camera


Note: In a pinhole camera we can adjust the focal length, all this will do is change the size of the resulting image

## Pinhole Camera (Simplified)

It is convenient to think of the image plane which is in from of the pinhole


## Pinhole Camera (Simplified)

It is convenient to think of the image plane which is in from of the pinhole


What happens if object moves towards the camera? Away from the camera?

## Focal Length

For a fixed sensor size, focal length determines the field of view (FoV)


## Focal Length

For a fixed sensor size, focal length determines the field of view (FoV)


Exercise: What is the field of view of a full frame ( 35 mm ) camera with a 50 mm lens? 100 mm lens?

## Focal Length



28 mm


50 mm


35 mm


70 mm

## Perspective Effects



Forsyth \& Ponce (2nd ed.) Figure 1.3a

## Perspective Effects

Far objects appear smaller than close ones


Forsyth \& Ponce (2nd ed.) Figure 1.3a

## Perspective Effects

Far objects appear smaller than close ones


Forsyth \& Ponce (2nd ed.) Figure 1.3a

## Perspective Effects

Far objects appear smaller than close ones


Forsyth \& Ponce (2nd ed.) Figure 1.3a
Size is inversely proportions to distance

## Perspective Effects

Far objects appear smaller than close ones


Forsyth \& Ponce (2nd ed.) Figure 1.3a

## Perspective Effects



Forsyth \& Ponce (1st ed.) Figure 1.3b

## Perspective Effects

Parallel lines meet at a point (vanishing point)


Forsyth \& Ponce (1st ed.) Figure 1.3b

## Vanishing Points

Each set of parallel lines meet at a different point

- the point is called vanishing point


## Vanishing Points

Each set of parallel lines meet at a different point

- the point is called vanishing point

Sets of parallel lines on the same plane lead to collinear vanishing points

- the line is called a horizon for that plane



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points

Each set of parallel lines meet at a different point

- the point is called vanishing point

Sets of parallel lines one the same plane lead to collinear vanishing points

- the line is called a horizon for that plane

Good way to spot fake images

- scale and perspective do not work
- vanishing points behave badly



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Vanishing Points



## Perspective Aside



## Perspective Aside



## Properties of Projection

- Points project to points
- Lines project to lines
- Planes project to the whole or half image
- Angles are not preserved


## Properties of Projection

- Points project to points
- Lines project to lines
- Planes project to the whole or half image
- Angles are not preserved


## Degenerate cases

- Line through focal point projects to a point
- Plane through focal point projects to a line


## Projection Illusion



## Projection Illusion



## Perspective Projection

3D object point


Forsyth \& Ponce (1st ed.) Figure 1.4

$$
P=\left[\begin{array}{l}
x \\
y \\
z
\end{array}\right] \text { projects to 2D image point } P^{\prime}=\left[\begin{array}{l}
x^{\prime} \\
y^{\prime}
\end{array}\right] \text { where }
$$

$$
\begin{aligned}
x^{\prime} & =f^{\prime} \frac{x}{z} \\
y^{\prime} & =f^{\prime} \frac{y}{z}
\end{aligned}
$$

Note: this assumes world coordinate frame at the optical center (pinhole) and aligned with the image plane, image coordinate frame aligned with the camera coordinate frame

## Perspective Projection: Proof



3D object point
Forsyth \& Ponce (1st ed.) Figure 1.4

$$
P=\left[\begin{array}{l}
x \\
y \\
z
\end{array}\right] \text { projects to 2D image point } P^{\prime}=\left[\begin{array}{c}
x^{\prime} \\
y^{\prime}
\end{array}\right] \text { where }
$$

$$
\begin{aligned}
& x^{\prime}=f^{\prime} \frac{x}{z} \\
& y^{\prime}=f^{\prime} \frac{y}{z}
\end{aligned}
$$

Note: this assumes world coordinate frame at the optical center (pinhole) and aligned with the image plane, image coordinate frame aligned with the camera coordinate frame

## Aside: Camera Matrix

Camera Matrix


Forsyth \& Ponce (1st ed.) Figure 1.4
$P=\left[\begin{array}{l}x \\ y \\ z \\ 1\end{array}\right]$ projects to 2D image point $P^{\prime}=\left[\begin{array}{c}x^{\prime} \\ y^{\prime} \\ 1\end{array}\right]$ where $P^{\prime}=\mathbf{C} P$

Note: this assumes world coordinate frame at the optical center (pinhole) and aligned with the image plane, image coordinate frame aligned with the camera coordinate frame

## Aside: Camera Matrix

Camera Matrix

$$
\begin{aligned}
& x^{\prime}=f^{\prime} \frac{x}{z} \\
& y^{\prime}=f^{\prime} \frac{y}{z}
\end{aligned}
$$

$$
\mathbf{C}=\left[\begin{array}{rrrr}
f^{\prime} & 0 & 0 & 0 \\
0 & f^{\prime} & 0 & 0 \\
0 & 0 & 1 & 0
\end{array}\right]
$$

$P=\left[\begin{array}{l}x \\ y \\ z \\ 1\end{array}\right]$ projects to 2D image point $P^{\prime}=\left[\begin{array}{c}x^{\prime} \\ y^{\prime} \\ 1\end{array}\right]$ where $P^{\prime}=\mathbf{C} P$

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$$
\mathbf{C}=\left[\begin{array}{rrrr}
f^{\prime} & 0 & 0 & 0 \\
0 & f^{\prime} & 0 & 0 \\
0 & 0 & 1 & 0
\end{array}\right]
$$

$$
\left[\begin{array}{rrrr}
f^{\prime} & 0 & 0 & 0 \\
0 & f^{\prime} & 0 & 0 \\
0 & 0 & 1 & 0
\end{array}\right]\left[\begin{array}{l}
x \\
y \\
z \\
1
\end{array}\right]=\left[\begin{array}{c}
f^{\prime} x \\
f^{\prime} y \\
z
\end{array}\right]=\left[\begin{array}{c}
\frac{f^{\prime} x}{f^{z}} \\
\frac{f^{\prime} y}{z} \\
1
\end{array}\right]
$$

$$
P=\left[\begin{array}{l}
x \\
y \\
z \\
1
\end{array}\right] \text { projects to 2D image point } P^{\prime}=\left[\begin{array}{c}
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$$

## Aside: Camera Matrix

Camera Matrix

$$
\mathbf{C}=\left[\begin{array}{rrrr}
f^{\prime} & 0 & 0 & 0 \\
0 & f^{\prime} & 0 & 0 \\
0 & 0 & 1 & 0
\end{array}\right]
$$

Pixels are squared / lens is perfectly symmetric
Sensor and pinhole perfectly aligned
Coordinate system centered at the pinhole


## Aside: Camera Matrix

Camera Matrix

$$
\mathbf{C}=\left[\begin{array}{rrrr}
f_{x}^{\prime} & 0 & 0 & 0 \\
0 & f_{y}^{\prime} & 0 & 0 \\
0 & 0 & 1 & 0
\end{array}\right]
$$



Sensor and pinhole perfectly aligned
Coordinate system centered at the pinhole


## Aside: Camera Matrix

Camera Matrix

$$
\mathbf{C}=\left[\begin{array}{rrrr}
f_{x}^{\prime} & 0 & 0 & c_{x} \\
0 & f_{y}^{\prime} & 0 & c_{y} \\
0 & 0 & 1 & 0
\end{array}\right]
$$



## Aside: Camera Matrix

Camera Matrix

$$
\mathbf{C}=\left[\begin{array}{rrrr}
f_{x}^{\prime} & 0 & 0 & c_{x} \\
0 & f_{y}^{\prime} & 0 & c_{y} \\
0 & 0 & 1 & 0
\end{array}\right] \mathbb{R}_{4 \times 4}
$$



$$
P=\left[\begin{array}{l}
x \\
y \\
z \\
1
\end{array}\right] \text { projects to 2D image point } P^{\prime}=\left[\begin{array}{c}
x^{\prime} \\
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1
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$$

## Aside: Camera Matrix

## Camera Matrix

$$
\mathbf{C}=\left[\begin{array}{rrrr}
f_{x}^{\prime} & 0 & 0 & c_{x} \\
0 & f_{y}^{\prime} & 0 & c_{y} \\
0 & 0 & 1 & 0
\end{array}\right] \mathbb{R}_{4 \times 4}
$$

Camera calibration is the process of estimating parameters of the camera matrix based on set of 3D-2D correspondences (usually requires a pattern whos structure and size is known)

$$
P=\left[\begin{array}{l}
x \\
y \\
z \\
1
\end{array}\right] \text { projects to 2D image point } P^{\prime}=\left[\begin{array}{c}
x^{\prime} \\
y^{\prime} \\
1
\end{array}\right] \text { where } P^{\prime}=\mathbf{C} P
$$

## Perspective Projection

3D object point


Forsyth \& Ponce (1st ed.) Figure 1.4

$$
P=\left[\begin{array}{l}
x \\
y \\
z
\end{array}\right] \text { projects to 2D image point } P^{\prime}=\left[\begin{array}{l}
x^{\prime} \\
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$$
\begin{aligned}
x^{\prime} & =f^{\prime} \frac{x}{z} \\
y^{\prime} & =f^{\prime} \frac{y}{z}
\end{aligned}
$$

Note: this assumes world coordinate frame at the optical center (pinhole) and aligned with the image plane, image coordinate frame aligned with the camera coordinate frame

## Weak Perspective



Forsyth \& Ponce (1st ed.) Figure 1.5

3D object point $P=\left[\begin{array}{l}x \\ y \\ z\end{array}\right]$ in $\Pi_{0}$ projects to 2D image point $P^{\prime}=\left[\begin{array}{c}x^{\prime} \\ y^{\prime}\end{array}\right]$
where $\begin{aligned} & \begin{array}{l}x^{\prime}=m x \\ y^{\prime}=m y\end{array}\end{aligned}$ and $m=\frac{f^{\prime}}{z_{0}}$

## Orthographic Projection



Forsyth \& Ponce (1st ed.) Figure 1.6

3D object point $P=\left[\begin{array}{l}x \\ y \\ z\end{array}\right]$
projects to 2D image point $P^{\prime}=\left[\begin{array}{l}x^{\prime} \\ y^{\prime}\end{array}\right]$
where $\left.\begin{array}{l}\left.\begin{array}{l}x^{\prime}= \\ y^{\prime}= \\ \end{array}\right]\end{array}\right]$

## Summary of Projection Equations

3D object point $P=\left[\begin{array}{l}x \\ y \\ z\end{array}\right]$ projects to 2D image point $P^{\prime}=\left[\begin{array}{c}x^{\prime} \\ y^{\prime}\end{array}\right]$ where

Perspective

$$
\begin{aligned}
& x^{\prime}=f^{\prime} \frac{x}{z} \\
& y^{\prime}=f^{\prime} \frac{y}{z} \\
& x^{\prime}=m x \quad m=\frac{f^{\prime}}{z_{0}} \\
& y^{\prime}=m y
\end{aligned}
$$

Orthographic

$$
x^{\prime}=x
$$

$$
y^{\prime}=y
$$

## Projection Models: Pros and Cons

Weak perspective (including orthographic) has simpler mathematics

- accurate when object is small and/or distant
- useful for recognition

Perspective is more accurate for real scenes

When maximum accuracy is required, it is necessary to model additional details of a particular camera

- use perspective projection with additional parameters (e.g., lens distortion)


## Why Not a Pinhole Camera?

- If pinhole is too big then many directions are averaged, blurring the image
- If pinhole is too small then diffraction becomes a factor, also blurring the image
- Generally, pinhole cameras are dark, because only a very small set of rays from a
 particular scene point hits the image plane
- Pinhole cameras are slow, because only a very small amount of light from a particular scene point hits the image plane per unit time



## Reason for Lenses

A real camera must have a finite aperture to get enough light, but this causes blur in the image


Solution: use a lens to focus light onto the image plane

## Reason for Lenses

A real camera must have a finite aperture to get enough light, but this causes blur in the image


The role of a lens is to capture more light while preserving, as much as possible, the abstraction of an ideal pinhole camera.


Solution: use a lens to focus light onto the image plane

## Snell's Law



$$
n_{1} \sin \alpha_{1}=n_{2} \sin \alpha_{2}
$$

## Snell's Law



$$
n_{1} \sin \alpha_{1}=n_{2} \sin \alpha_{2}
$$

## Pinhole Model (Simplified) with Lens



## General Lens

Ray of light

## Thin Lens


https://phys.libretexts.org/Bookshelves/University Physics/Book\%3A University Physics_(OpenStax)/Map\%3A University Physics_III__Optics_and_Modern_Physics_(OpenStax)/02\%3A_Geometric_Optics_and_Image_Formation/2.05\%3A_Thin_Lenses

## Thin Lens Equation



Forsyth \& Ponce (1st ed.) Figure 1.9

$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$

## Thin Lens Equation



Forsyth \& Ponce (1st ed.) Figure 1.9

$$
\frac{1}{z^{\prime}}+\frac{1}{z}=\frac{1}{f}
$$

## Thin Lens Equation



Forsyth \& Ponce (1st ed.) Figure 1.9

$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$

## Thin Lens Equation: Derivation



Forsyth \& Ponce (1st ed.) Figure 1.9

$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$

## Thin Lens Equation: Derivation

$$
\begin{array}{|l|l|}
\hline \frac{y}{-z}-\frac{-y^{\prime}}{-z^{\prime}} \\
\frac{y}{y^{\prime}}=\frac{z}{z^{\prime}}
\end{array}
$$



Forsyth \& Ponce (1st ed.) Figure 1.9

$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$

## Thin Lens Equation: Derivation



$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$

## Thin Lens Equation: Derivation



## Thin Lens Equation: Derivation



## Possible Uses of Thin Lens Abstraction



Forsyth \& Ponce (1st ed.) Figure 1.9

$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$

## Lens Basics

A lens focuses parallel rays (from points at infinity) at focal length of the lens
Rays passing through the center of the lens are not bent


## Lens Basics

Lenses focus all rays from a (parallel to lense) plane in the world

$$
\frac{1}{z^{\prime}}-\frac{1}{z}=\frac{1}{f}
$$



Objects off the plane are blurred depending on the distance

## Effect of Aperture Size



Smaller aperture $\Rightarrow$ smaller blur, larger depth of field


## Depth of Field



Aperture size $=\mathrm{f} / \mathrm{N}, \Rightarrow$ large $\mathrm{N}=$ small aperture

