Notes

- Errors in last lecture missing density in viscosity terms:
 - Incompressible Navier-Stokes is

$$u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g + \frac{1}{\rho} \nabla \cdot \mu \Big(\nabla u + \nabla u^T \Big)$$

 $\nabla \cdot \boldsymbol{u} = \boldsymbol{0}$

· With constant viscosity, momentum equation is

$$u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g + \frac{\mu}{\rho} \nabla^2 u$$

- Often (particularly if density is constant) take parameter $\nu{=}\mu{/}\rho$ to get

$$u_t + u \cdot \nabla u + \frac{1}{\rho} \nabla p = g + v \nabla^2 u$$

Nondimensionalization

- Actually go even further
- Select a characteristic length L
 - e.g. the width of the domain,
- · And a typical velocity U
 - e.g. the speed of the incoming flow
- Rescale terms
 - x'=x/L, u'=u/U, t'=tU/L, p'=p/ ρ U² so they all are dimensionless $u'_t + u' \cdot \nabla u' + \nabla p' = \frac{Lg}{U^2} + \frac{v}{UL} \nabla^2 u'$

Nondimensional parameters

- Re=UL/v is the Reynold's number
 - The smaller it is, the more viscosity plays a role in the flow
 - High Reynold's numbers are hard to simulate
- $\operatorname{Fr} = U / \sqrt{|g|L}$ is the Froude number
 - The smaller it is, the more gravity plays a role in the flow
 - Note: often can ignore gravity (pressure gradient cancels it out)

Vorticity

- Last class: irrotational flow
 - And simplification to potential flow
- How do we measure rotation?
 - Vorticity of a vector field (velocity) is: $\omega = \nabla \times u$
 - Proportional (but not equal) to angular velocity of a rigid body off by a factor of 2
- · Visualization of potential flow is fairly boring
 - It's as smooth as possible, laminar
 - · Vorticity is what makes flow look cool
 - (Or free surfaces...)

Vorticity equation

• Start with N-S, constant viscosity and density

 $u_t + u \cdot \nabla u + \frac{1}{2} \nabla p = g + v \nabla^2 u$

- Take curl of whole equation $\nabla \times u_t + \nabla \times (u \cdot \nabla u) + \nabla \times \frac{1}{\rho} \nabla p = \nabla \times g + \nabla \times (v \nabla^2 u)$
- Lots of terms are zero:
 - g is constant (or the potential of some field)
 - With constant density, pressure term too $\nabla \times u_t + \nabla \times (u \cdot \nabla u) = v \nabla \times \nabla^2 u$
- Then use vector identities to simplify... $\nabla \times u_t + \nabla \times ((\nabla \times u) \times u + \frac{1}{2} \nabla u^2) = v \nabla^2 (\nabla \times u)$ $\omega_t + \nabla \times (\omega \times u) = v \nabla^2 \omega$

Vorticity equation continued

• Simplify with more vector identities, and assume incompressible to get:

$$\frac{D\omega}{Dt} = \omega \cdot \nabla u + v \nabla^2 \omega$$

- Important result: Kelvin Circulation Theorem
 - Roughly speaking: if ω=0 initially, and there's no viscosisty, ω=0 forever after (following a chunk of fluid)
- If fluid starts off irrotational, it will stay that way (in many circumstances)
- So potential flow is reasonable

Potential in time

- Use vector identity u•∇u=(∇×u)×u+∇lul²/2
- Assume
 - incompressible (∇•u=0), inviscid, irrotational (∇×u=0)
 - · constant density
 - thus potential flow (u= $\nabla \phi$, $\nabla^2 \phi$ =0)
- Then momentum equation simplifies (using G=-gy for gravitational potential)

$$u_t + (\nabla \times u) \times u + \frac{1}{2} \nabla |u|^2 + \frac{1}{\rho} \nabla p = g$$
$$\nabla \phi_t + \frac{1}{2} \nabla |u|^2 + \frac{1}{\rho} \nabla p = -\nabla G$$

Bernoulli's equation

• Every term in the simplified momentum equation is a gradient: integrate to get

$$\phi_t + \frac{1}{2}u^2 + \frac{p}{\rho} = -G$$

- (Remember Bernoulli's law for pressure?)
- This tells us how the potential should evolve in time

Water waves

- For small waves (no breaking), can reduce geometry of water to 2D heightfield
- Can reduce the physics to 2D also
 - How do surface waves propagate?
- Plan of attack
 - Start with potential flow, Bernoulli's equation
 - Write down boundary conditions at water surface
 - Simplify 3D structure to 2D

Set up

- We'll take y=0 as the height of the water at rest
- H is the depth (y=-H is the sea bottom)
- h is the current height of the water at (x,z)
- Simplification: velocities very small (small amplitude waves)

Boundaries

- At sea floor (y=-H), v=0 $\phi_y = 0$
- At sea surface (y=h≈0), v=h_t
 - Note again assuming very small horizontal motion $\phi_v = h_t$
- At sea surface (y=h≈0), p=0
 - Or atmospheric pressure, but we only care about pressure differences
 - Use Bernoulli's equation, throw out small velocity squared term, use p=0,

$$\phi_t = -gh$$

Finding a wave solution

- Plug in φ=f(y)sin(K•(x,z)-ωt)
 - In other words, do a Fourier analysis in horizontal component, assume nothing much happens in vertical
 - Solving $\nabla^2 \phi = 0$ with boundary conditions on ϕ_y gives $\phi = A \frac{\omega}{|K|} \frac{\cosh(|K|(y+H))}{\sinh(|K|H)} \sin(K \cdot (x,z) - \omega t)$
 - Pressure boundary condition then gives (with k the magnitude of K)

$$\omega = \sqrt{gk \tanh kH}$$

Dispersion relation

• So the wave speed c is

$$c = \frac{\omega}{k} = \sqrt{\frac{g}{k}} \tanh kH$$

- Notice that waves of different wavenumbers k have different speeds
 - Separate or disperse in time
- For deep water (H big, k reasonable -not tidal waves!) tanh(kH)≈1

$$c = \sqrt{\frac{g}{k}}$$

Simulating the ocean

- So far from land, a reasonable thing to do is
 - Do Fourier decomposition of initial surface height
 - Evolve each wave according to given wave speed (dispersion relation)
 - Update phase, use FFT to evaluate
- How do we get the initial spectrum?
 - Measure it! (oceanography)

Energy spectrum

• Fourier decomposition of height field:

$$h(x,z,t) = \sum_{i,j} \hat{h}(i,j,t) e^{\sqrt{-1}(i,j)\cdot(x,z)}$$

- "Energy" in K=(i,j) is $S(K) = |\hat{h}(K)|^2$
- Oceanographic measurements have found models for expected value of S(K) (statistical description)

Phillips Spectrum

- For a "fully developed" sea
 - wind has been blowing a long time over a large area, statistical distribution of spectrum has stabilized
- The Phillips spectrum is: [Tessendorf...]

$$S(K) = A \frac{1}{k^4} \exp\left(\frac{-1}{(kL)^2} - (kl)^2\right) \left(\frac{|K \cdot W|}{|K||W|}\right)^2$$

- A is an arbitrary amplitude
- L=IWI²/g is largest size of waves due to wind velocity W and gravity g
- Little I is the smallest length scale you want to model

Other spectra

- More complex models such as JONSWAP
 - Sea is never fully developed, need to take into account how far from land you are
- Or make up your own

Fourier synthesis

• From the prescribed S(K), generate actual Fourier coefficients

$$\hat{h}(K,0) = \frac{1}{\sqrt{2}} (X_1 + X_2 \sqrt{-1}) \sqrt{S(K)}$$

- X_i is a random number with mean 0, standard deviation 1 (Gaussian)
- Uniform numbers from unit circles aren't terrible either
- Want real-valued h, so must have $\hat{h}(K) = \hat{h}(-K)^*$
 - Or give only half the coefficients to FFT routine and specify you want real output

Time evolution

- Dispersion relation gives us $\omega(K)$
- At time t, want $h(x,t) = \sum_{K=(i,j)} \hat{h}(K,0)e^{\sqrt{-1}(K \cdot x \omega t)}$

$$= \sum_{K-(i,i)} \hat{h}(K,0) e^{-\sqrt{-1}\omega t} e^{\sqrt{-1}K \cdot x}$$

- So then coefficients at time t are
 - For j≥0: $\hat{h}(i, j, t) = \hat{h}(i, j, 0)e^{-\sqrt{-1}\omega t}$
 - Others: figure out from conjugacy condition (or leave it up to real-valued FFT to fill them in)

Picking parameters

- Need to fix grid for Fourier synthesis (e.g. 1024x1024 height field grid)
- Grid spacing shouldn't be less than e.g. 2cm (smaller than that - surface tension, nonlinear wave terms, etc. take over)
 - Take little I (cut-off) a few times larger
- Total grid size should be greater than but still comparable to L in Phillips spectrum (depends on wind speed and gravity)
- Amplitude A shouldn't be too large
 - Assumed waves weren't very steep

Note on FFT output

- FFT takes grid of coefficients, outputs grid of heights
- It's up to you to map that grid (0...n-1, 0...n-1) to world-space coordinates
- In practice: scale by something like L/n
 - Adjust scale factor, amplitude, etc. until it looks nice
- Alternatively: look up exactly what your FFT routines computes, figure out the "true" scale factor to get world-space coordinates

Tiling issues

- Resulting grid of waves can be tiled in x and z
- Handy, except people will notice if they can see more than a couple of tiles
- Simple trick: add a second grid with a nonrational multiple of the size
 - Golden mean (1+sqrt(5))/2=1.61803... works well [why?]
 - The sum is no longer periodic, but still can be evaluated anywhere in space and time easily enough

Choppy waves

- See Tessendorf for more explanation
- Nonlinearities cause real waves to have sharper peaks and flatter troughs than linear Fourier synthesis gives
- Can manipulate height field to give this effect
 - Distort grid with $(x,z) \rightarrow (x,z)+\lambda D(x,z,t)$

$$D(x,t) = \sum_{K} -\sqrt{-1} \frac{K}{|K|} \hat{h}(K,t) e^{\sqrt{-1}K}$$

Choppiness problems

- The distorted grid can actually tangle up (Jacobian has negative determinant not 1-1 anymore)
 - Can detect this, do stuff (add particles for foam, spray?)
- Can't easily use superposition of two grids to defeat periodicity... (but with a big enough grid and camera position chosen well, not an issue)

Issues with Fourier method

- Can't easily handle objects in water
 - E.g. boat wakes, splashes
- One solution: cover up problem with foam
- While dispersion relation works in shallow water too, can't handle beaches...
- Next class: shallow water equations (PDE's)