Notes on Last Lecture

- Approximate interior normal may be quite wrong [corner example]
- Lots of potential ways to fix this if it happens
 - Fall back on collision detection (normal at collision point on surface should work)
 - If the object normal doesn't work, use opposite of particle velocity instead (maybe too inelastic)
 - Use a repulsion impulse (and friction) to get out: $\Delta v_N = (-\phi/\Delta t)n$ (dangerous: adds energy!)
- Unfortunately, to be robust enough, usually need to throw in a bunch of hacks...

Notes

- Round-off error is also a problem
- In algorithms described before, can get into infinite loops if not careful: v_N^{before}+Δv_N ≠ v_N^{after}
 - If collision resolution doesn't seem to ever converge---could just be round-off
 - Simple fix: stop after a fixed number of iterations, keep the old particle position
 - Not so easy with moving objects really need to update position
 - So use very weak repulsions to push objects just slightly clear of objects

Moving triangles

- · Life is a little more complicated
- Assume corners of the triangles move in linear trajectories too
 - Note this is NOT rigid in general...
- At time s, corner is at x_i+sv_i
 - (assume s starts at 0 at start of time step)
- Normal is also changing in time
 - So for plane intersection, need to substitute for n the cross-product formula
 - Thankfully, don't need to normalize, since that doesn't change the plane

Moving triangles equation

• A cubic in s to solve:

 $\left[(1-s)p + sq - x_1 \right] \bullet \left[\left(x_2(s) - x_1(s) \right) \times \left(x_3(s) - x_1(s) \right) \right] = 0$

- Only interested in real solutions between 0 and Δt
- · Solve iteratively
 - Derivative=quadratic can be solved to tell us if any extrema in interval
 - Values at endpoints and at any extrema in interval tell us the intervals that roots could be in
 - Solve for those roots with secant/bisection search

Acceleration

- Too slow to check every single triangle if mesh is large
- Need acceleration
- Also critical if we have lots of distinct objects (even if implicit)
- Lots of papers written on acceleration structures
 - Prune out unnecessary tests

Bounding Volume Hierarchy

- Surround each triangle (or small group of triangles) with a simple bounding volume
 - E.g. axis-aligned box, sphere, oriented box...
- Surround group of bounding volumes with a parent bounding volume, and so on up
- End up with a tree
- To check a segment against scene, check if it could overlap root of tree
 - If not, we're done
 - If so, recurse on children

Grid Acceleration

- Or, put down a virtual grid in space
 - Each grid cell has a list of which triangles overlap
- To test a segment, only look at triangles in the grid cells the segment crosses
- · Can use hash table for memory efficiency
 - Hash on cell indices (i,j,k)
- Note trade-off:
 - The finer the grid, the fewer extraneous triangles
 - But: more grid cells to check, more memory used, and more expensive to build grid
 - Tune for your application!

Rigid Bodies

- Very well studied
- I'll introduce them from a particle perspective
 - · Easy to get lost in abstract notions
 - Particles are fundamental
- Discretize an object into small point masses
 - x_i, v_i, m_i
- Assume object doesn't change shape (doesn't deform)
 - What does that mean for the motion of the particles? How do we describe it, solve for it?

World Space vs. Object Space

- World space: where the particles actually are now
 - This is where we will look at x, v, and almost every other quantity
- Object space: imaginary "reference" place for the particles
 - Label the object space position p_i
 - Does not change as the object moves things we compute in object space stay constant
 - We can define it arbitrarily

Mapping

- The map from p_i to x_i(t) cannot change the shape
 - The distance between any two particles never changes
 - Thus map has to be x_i(t)=R(t)p_i+X(t)
 - R(t) is an orthogonal 3x3 matrix: RR^T= δ
 - The orientation (rotation) of the object
 - X(t) is a vector
 - The "location" of the object

Rigid Motion

- Differentiate map w.r.t. time (using dot notation): v_i = Rp_i + V
- Invert map for p_i : $p_i = R^T(x_i X)$
- Thus: $v_i = \dot{R}R^T(x_i X) + V$
- 1st term: rotation, 2nd term: translation
 Let's simplify the rotation

Skew-Symmetry

• Differentiate $RR^T = \delta$ w.r.t. time:

$$\dot{R}R^{T} + R\dot{R}^{T} = 0 \implies \dot{R}R^{T} = -(\dot{R}R^{T})^{T}$$

• Skew-symmetric! Thus can write as:

$$\dot{R}R^{T} = \begin{pmatrix} 0 & -\omega_{2} & \omega_{1} \\ \omega_{2} & 0 & -\omega_{0} \\ -\omega_{1} & \omega_{0} & 0 \end{pmatrix}$$

• Call this matrix ω^* (built from a vector ω) $\dot{R}R^T = \omega^* \implies \dot{R} = \omega^* R$

The cross-product matrix

• Note that:

$$\omega^* x = \begin{pmatrix} 0 & -\omega_2 & \omega_1 \\ \omega_2 & 0 & -\omega_0 \\ -\omega_1 & \omega_0 & 0 \end{pmatrix} \begin{pmatrix} x_0 \\ x_1 \\ x_2 \end{pmatrix} = \begin{pmatrix} \omega_1 x_2 - \omega_2 x_1 \\ \omega_2 x_0 - \omega_0 x_2 \\ \omega_0 x_1 - \omega_1 x_0 \end{pmatrix} = \omega \times x$$

• So we have:

$$v_i = \omega \times (x_i - X) + V$$

- $\boldsymbol{\omega}$ is the angular velocity of the object

Angular velocity

- Recall:
 - $\ensuremath{\omega}\xspace$ l $\ensuremath{\omega}\xspacee$ l $\ensuremath{\omega}\xspace$
 - ω points along the axis of rotation (which in this case passes through the point X)
 - Convince yourself this makes sense with the properties of the cross-product

Force

- Take another time derivative to get acceleration: $a_i = \dot{v}_i = \ddot{R}p_i + A$
- Use F=ma, sum up net force on system: $\sum_{i} F_{i} = \sum_{i} m_{i} a_{i} = \sum_{i} m_{i} (\ddot{R}p_{i} + A)$ $= \ddot{R} \sum_{i} m_{i} p_{i} + A \sum_{i} m_{i}$
- Let the total mass be $M = \sum_{i} m_{i}$
- How to simplify the other term?

Centre of Mass

• Let's pick a new object space position:

$$p_i^{new} = p_i - \frac{\sum_j m_j p_j}{M}$$

- The mass-weighted average of the positions is the centre of mass
- We translated the centre of mass (in object space) to the point 0
- Now: $\sum_i m_i p_i = 0$

Force equation

- So now, assuming we've set up object space right (centre of mass at 0), F=MA
- If there are no external forces, have F=0
 - Internal forces must balance out, opposite and equal
 - Thus A=0, thus V=constant
- If there are external forces, can integrate position of object just like a regular particle!

What about R?

- How does orientation change?
- Think about internal forces keeping the particles in the rigid configuration
 - Conceptual model: very stiff spring between every pair of particles, maintaining the rest length
- So $F_i = \sum_j f_{ij}$ where f_{ij} is force on i due to j
- Of course f_{ii}+f_{ii}=0
- Also: f_{ii} is in the direction of x_i-x_j
 - Thus $(x_i x_j) \times f_{ij} = 0$

Net Torque

- Play around: $\frac{((x_i X) (x_j X)) \times f_{ij} = 0}{(x_i X) \times f_{ij}} = (x_j X) \times f_{ij}}$ $= -(x_j X) \times f_{ij}$
- Sum both sides (look for net force)

$$\sum_{i,j} (x_i - X) \times f_{ij} = -\sum_{i,j} (x_j - X) \times f_{ji}$$
$$\sum_i (x_i - X) \times F_i = -\sum_j (x_j - X) \times F_j$$
$$= 0$$

• The expression we just computed=0 is the net torque on the object

Torque

• The torque of a force applied to a point is

$$\boldsymbol{\tau}_i = (\boldsymbol{x}_i - \boldsymbol{X}) \times \boldsymbol{F}_i$$

- The net torque due to internal forces is 0
- [geometry of torque: at CM, with opposite equal force elsewhere]
- Torque obviously has something to do with rotation
- How do we get formula for change in angular velocity?

Angular Momentum

• Use F=ma in definition of torque:

$$\tau_i = (x_i - X) \times m_i a_i$$
$$= \frac{d}{dt} \Big[m_i (x_i - X) \times v_i \Big]$$

- force=rate of change of linear momentum, torque=rate of change of angular momentum
- The total angular momentum of the object is

$$L = \sum_{i} m_{i}(x_{i} - X) \times v_{i}$$
$$= \sum_{i} m_{i}(x_{i} - X) \times (v_{i} - V)$$

Getting to ω

- Recall $v_i V = \omega \times (x_i X)$
- Plug this into angular momentum:

$$L = \sum_{i} m_{i}(x_{i} - X) \times (\omega \times (x_{i} - X))$$

= $-\sum_{i} m_{i}(x_{i} - X) \times ((x_{i} - X) \times \omega)$
= $-\sum_{i} m_{i}(x_{i} - X)^{*}(x_{i} - X)^{*}\omega$
= $\underbrace{\left(\sum_{i} m_{i}(x_{i} - X)^{*T}(x_{i} - X)^{*}\right)}_{I(t)}\omega$

Inertia Tensor

- I(t) is the inertia tensor
- Kind of like "angular mass"
- Linear momentum is mv
- Angular momentum is L=I(t)ω
- Or we can go the other way: $\omega = I(t)^{-1}L$

Equations of Motion

$$\frac{\frac{d}{dt}V = F/M \quad \frac{d}{dt}L = T$$

$$\frac{\frac{d}{dt}X = V \qquad \omega = I(t)^{-1}L$$

$$\frac{\frac{d}{dt}R = \omega^*R$$

In the absence of external forces F=0, T=0

Reminder

- Before going on:
- · Remember that this all boils down to particles
 - Mass, position, velocity, (linear) momentum, force are fundamental
 - Inertia tensor, orientation, angular velocity, angular momentum, torque are just abstractions
 - Don't get too puzzled about interpretation of torque for example: it's just a mathematical convenience

Inertia Tensor Simplified

• Reduce expense of calculating I(t):

$$I(t) = \sum_{i} m_{i} (x_{i} - X)^{*T} (x_{i} - X)^{*}$$

=
$$\sum_{i} m_{i} [(x_{i} - X)^{T} (x_{i} - X) \delta - (x_{i} - X) (x_{i} - X)^{T}]$$

- Now use $x_i\text{-}X\text{=}Rp_i$ and use $R^{\mathsf{T}}R\text{=}\delta$

$$I(t) = \sum_{i} m_{i} \left[p_{i}^{T} R^{T} R p_{i} \delta - R p_{i} p_{i}^{T} R^{T} \right]$$
$$= R \underbrace{\left(\sum_{i} m_{i} \left(p_{i}^{T} p_{i} \delta - p_{i} p_{i}^{T} \right) \right)}_{I_{body}} R^{T}$$

Inertia Tensor Simplified 2

- So just compute inertia tensor once, for object space configuration
- Then I(t)=RI_{body}R^T
- And $I(t)=R(I_{body})^{-1}R^{T}$
 - So precompute inverse too
- In fact, since I is symmetric, know we have an orthogonal eigenbasis Q
- Rotate object-space orientation by Q
 - Then I_{body} is just diagonal!

Degenerate Inertia Tensors

- I is just sum of symmetric positive semidefinite matrices
 - Each one has null space: vectors parallel to x_i-X
- If all the points line up (object is a rod) then sum I has the same null space
 - Singular: cannot be inverted
 - We don't care though, since we can't track rotation around that axis anyways
 - · So diagonalize I, and only invert nonzero elements
- Similarly for a single point...

Taking the limit

- Letting our decomposition of the object into point masses go to infinity:
 - Instead of sum over particles, integral over object volume
 - Instead of particle mass, density at that point in space

$$\sum_{i} m_i \operatorname{foo}(x_i) \to \iiint_{x} \rho(x) \operatorname{foo}(x) dx$$

• No big deal

Approximating Inertia Tensors

- For complicated geometry, don't really need exact answer
- · Instead use numerical quadrature
 - If we can afford to spend a lot of time precomputing, life is simple
 - Simplest approach: Monte-Carlo
 - Obviously stratified sampling etc. helps

Computing Inertia Tensors

- Do the integrals: $I_{body} = \iiint \rho (p^T p \delta p p^T) dp$
- Lots of fun!
- You may want to look them up instead
 - E.g. Eric Weisstein's World of Science on the web
- Align axis perpendicular to planes of symmetry (of ρ) in object space
 - Guarantees some off-diagonal zeros
- Example: sphere, uniform density, radius R



Combining Objects

- · What if object is union of two simpler objects?
- Integrals are additive
 - But be careful about adding I₁(t)+I₂(t):
 - World-space formulas (x-X) use the X for the object: X₁ and X₂ may be different
 - Simplified $\mathbf{I}_{\mathrm{body}}$ formula based on having centre of mass at origin
 - Let's work it out from the integral of I(t)
- Combined mass: M=M₁+M₂
- Centre of mass of combined object:

$$X = \frac{\int_{\Omega_{1} \cup \Omega_{2}} \rho x}{\int_{\Omega_{1} \cup \Omega_{2}} \rho} = \frac{M_{1}X_{1} + M_{2}X_{2}}{M}$$

Combined Inertia Tensor

$$\begin{split} I(t) &= \int_{\Omega_1 \cup \Omega_2} \rho(x - X)^{*T} (x - X)^* \\ &= \int_{\Omega_1} \rho(x - X_1 + X_1 - X)^{*T} (x - X_1 + X_1 - X)^* + \int_{\Omega_2} \cdots \\ &= \int_{\Omega_1} \rho(x - X_1)^{*T} (x - X_1)^* + \int_{\Omega_1} \rho(X_1 - X)^{*T} (x - X_1)^* \\ &+ \int_{\Omega_1} \rho(x - X_1)^{*T} (X_1 - X)^* + \int_{\Omega_1} \rho(X_1 - X)^{*T} (X_1 - X)^* + \int_{\Omega_2} \cdots \\ &= I_1(t) + (X_1 - X)^{*T} \underbrace{\int_{\Omega_1} \rho(x - X_1)^*}_{0} + \underbrace{\int_{\Omega_1} \rho(x - X_1)^{*T}}_{0} (X_1 - X)^{*T} (X_1 - X)^* \\ &+ M_1(X_1 - X)^{*T} (X_1 - X)^* + \int_{\Omega_2} \cdots \\ &= I_1(t) + M_1(X_1 - X)^{*T} (X_1 - X)^* + I_2(t) + M_2(X_2 - X)^{*T} (X_2 - X)^* \end{split}$$

Numerical Method

- For advancing V and X, can use any of the second order schemes we discussed before
 - Often only gravity and small amount of wind drag
- For advancing angular stuff:
 - Constraint on R makes life a little more interesting

Advancing angular stuff

• Symplectic Euler-like algorithm simplest choice: $L_{n+1} = L_n + \Delta t T$

$$\omega_{n+1} = I(t_n)^{-1}L_{n+1}$$
$$R_{n+1} = R_n + \Delta t \omega_{n+1}^* R_n$$

- Note: updated R isn't quite orthogonal
- Need to correct (otherwise objects inflate)
- Simplest choice: Gram-Schmidt
 - But introduces axis-bias, and expensive
- Could also compute rotation matrix for $\Delta t \omega$
 - · Even more expensive, still have some drift

Stability? Accuracy?

- Note R cannot blow up (we keep making it orthogonal)
- But if T=T(R,ω) there is potential for L and ω to blow up
 - Rarely the case (usually T=0, apart from isolated collision impulses)
 - If it is the case, can go implicit
- May want to restrict Δt=O(ω⁻¹) to properly sample rotations

Improving on R

- Expensive (and maybe biased) to keep R orthogonal
 - 9 numbers for 3 parameters
 - · Use a less redundant representation
- Quaternions work better!
 - · Still cheap and easy to deal with (unlike Euler angles, for example)
 - · Only 4 numbers still need to normalize
 - But can do it without axis bias
 - and for much cheaper

Review quaternions

- Instead of R, use q=(s,x,y,z) with lql=1
 - Can think of q=s+xi+yj+zk
 - i²=j²=k²=1, ij=-ji=k, jk=-kj=i, ki=-ik=j
 - Don't commute! $q_1q_2 \neq q_2q_1$
- Represents "half" a rotation:
 - q=cos(θ/2)
 - $|x,y,z|^{2}=\sin^{2}(\theta/2)$
 - Axis of rotation is (x,y,z)
- Conjugate (inverse for unit norm) is

 $\overline{q} = (s, -x, -y, -z)$

Rotating with quaternions

- Instead of Rp, calculate $q(0,p)\overline{q}$
- Composing a rotation of Δtω to advance a $q_{n+1} = \left(\sqrt{1 - \left|\Delta t \frac{\omega}{2}\right|^2}, \Delta t \frac{\omega}{2}\right) q_n$ time step:
- For small Δtω approximate:

$$q_{n+1} = \left(1, \Delta t \frac{\omega}{2}\right) q_n = q_n + \Delta t \frac{\omega}{2} q_n$$

From this get the differential equation:

$$\dot{q} = \frac{1}{2}\omega q$$

Converting q to R

- Clearly superior to use guaternions for storing and updating orientation
- · But, slightly faster to transform points with rotation matrix
- If you need to transform a lot of points (collision detection...) may want to convert q into R
- Basic idea: columns of R are rotated axes R(1,0,0)^T, R(0,1,0)^T, and R(0,0,1)^T
- Do the rotation with g instead.
 - · Can simplify and optimize for the zeros look it up